

# Strangeness Report

Federico Antinori INFN Padova - Italy



#### Contents

- → I concentrate on strange particle abundances (yields and ratios)
- QGP predictions
- Enhancement pattern
- Hadronic transport
- Thermal models
- Beyond equilibrium?
- The  $\Phi$
- Extras
- Conclusions

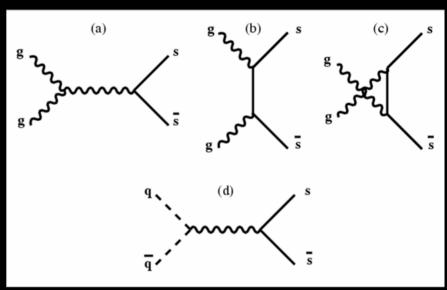
not covered: Resonances (Fachini), Spectra (Velkovska), Flow (Retiere), Pentaguarks (Jaffe), HBT (Magestro)

## Two historic QGP predictions

- $\bullet$  restoration of  $\chi$  symmetry  $\rightarrow$  increased production of s
  - mass of strange quark in QGP expected to go back to current value
    - m<sub>s</sub> ~ 150 MeV ~ Tc
  - → copious production of s̄s pairs, mostly by gg fusion

[Rafelski: Phys. Rep. 88 (1982) 331]

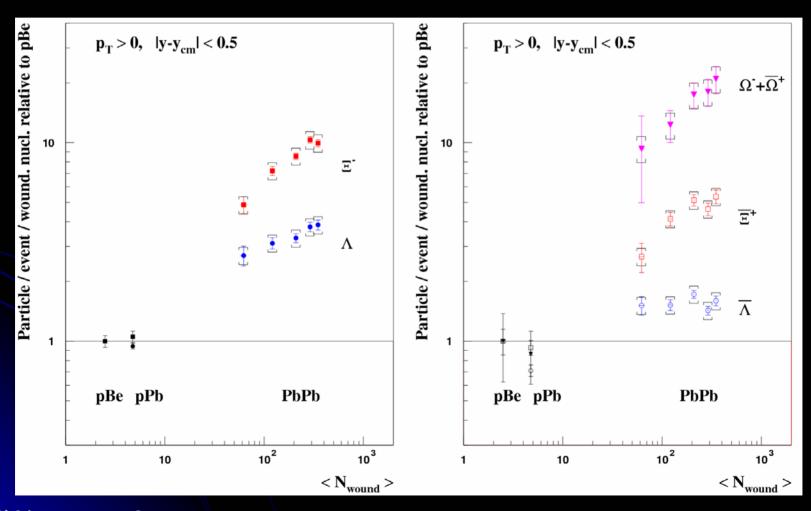
[Rafelski-Müller: P. R. Lett. 48 (1982) 1066]



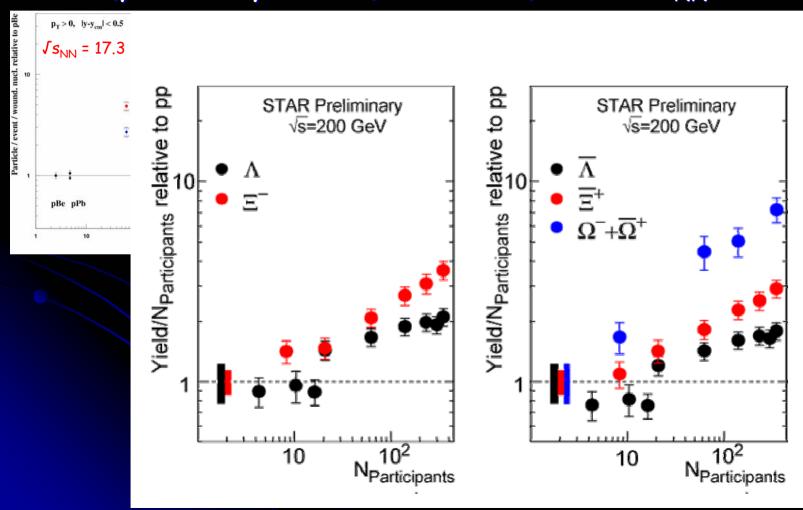
- deconfinement → stronger effect for multi-strange
  - can be built using uncorrelated s quarks produced in independent microscopic reactions
  - strangeness enhancement increasing with strangeness content

[Koch, Müller & Rafelski: Phys. Rep. 142 (1986) 167]

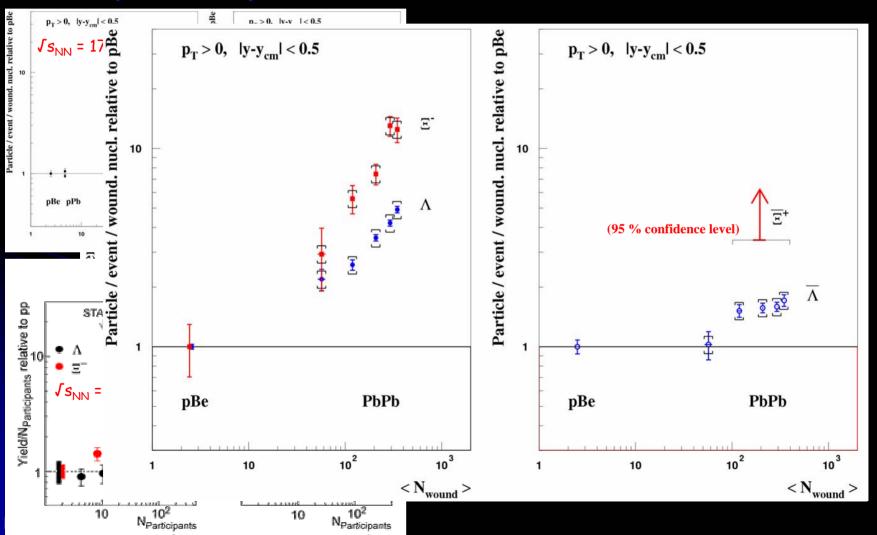
• @ top SPS (NA57,  $\int s_{NN} = 17.3 \, GeV$ )

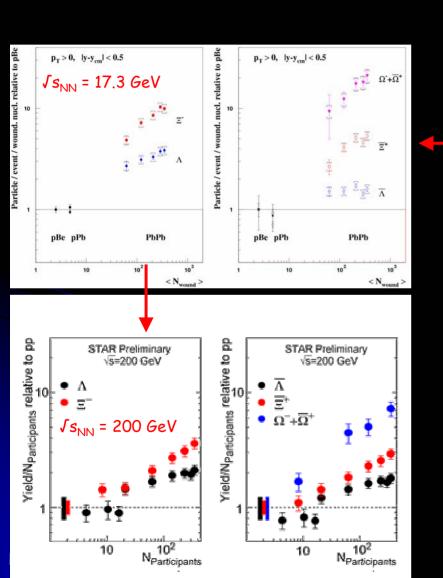


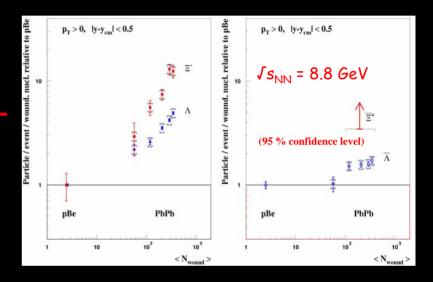
and (yesterday's news) @ RHIC (STAR, \( \sigma\_{\text{NN}} = 200 \text{ GeV} \)



and (yesterday's news) @ low E SPS (NA57, \( \sum\_{NN} = 8.8 \) GeV)







 qualitatively similar picture emerging from 8.8 to 200 GeV

#### Alternative definition of pA enhancement?

- e.g.: [Fischer: Nucl. Phys. A 715 (2003) 118]
- two-component model of particle production:

⇒ in pA: 
$$Y_{pA} = [v/2 + 1/2] Y_{pp}$$
 (v = number of collisions)  
=  $N_{part} Y_{pp}/2$ 

enhancement "blamed" on projectile component only:
 usual enhancement: E alternative enhancement: F

$$(Y/N_{part})_{pA} = E \cdot (Y/N_{part})_{pp}$$
  
 $\rightarrow Y_{pA} = E [v/2 + 1/2] Y_{pp}$ 

$$Y_{pA} = [v/2 + F/2] Y_{pp}$$

• then:  $(F - 1) = (E - 1) N_{part}$ 

reasonable for baryon number transfer, but do e.g. strange quarks know if they come from projectile or target?

 $\rightarrow$  enhancement within pA appears even if pA yields  $\propto$   $N_{part}$ 

#### Hadronic transport

- Hadronic transport codes:
  - do reasonably well on singly strange particles
  - but fail to reproduce the production of multi-strange particles at SPS and RHIC
    - see for instance:
      - [Soff et al.: Phys. Lett. B471 (1999) 89],
      - [C.Greiner: nucl-th/0208080 and references there],
      - [STAR: nucl-ex/0307024],
      - [Huovinen & Kapusta: nucl-th/0310051]

[Huovinen & Kapusta: nucl-th/0310051]

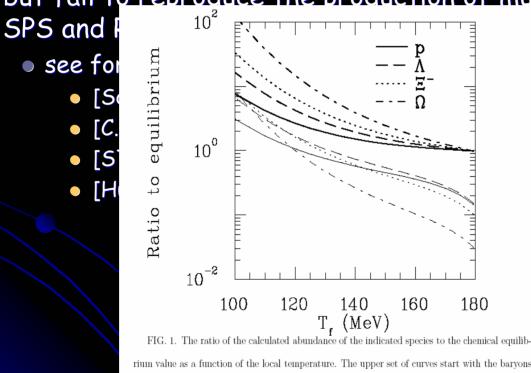
## Hadronic transport

Hadronic transport codes:

do reasonably well on singly strange particles

but fail to reproduce the production of multi-strange particles at

e],



rium value as a function of the local temperature. The upper set of curves start with the baryons in equilibrium at  $T_c$ , the lower set start with no baryons.

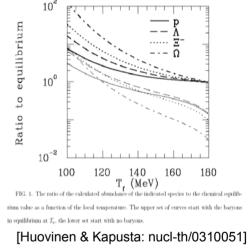
[Huovinen & Kapusta: nucl-th/0310051]

#### Hadronic transport

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  - do reasonably well on singly strange particles

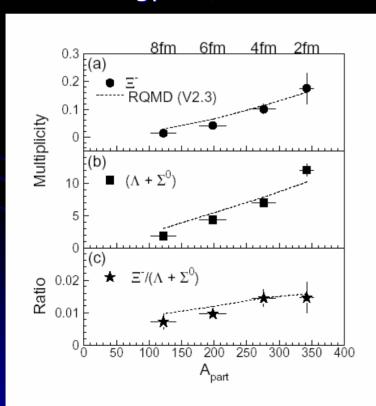
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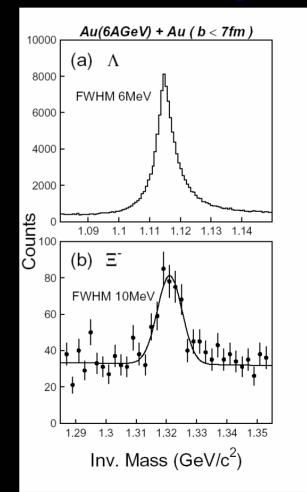
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  - [Huovinen & Kapusta: nucl-th/0310051]
- they get closer if:
  - masses are reduced towards chiral values or string tension is enhanced with respect to ee/pp/pA
  - enhanced contribution from inelastic scattering during expansion
- Activity is still ongoing:
  - e.g.: [Capella: nucl-th/0303045; Subrata Pal et al.: nucl-th/0106074]
  - hadronic models are getting more and more exotic...



#### Meanwhile at the AGS...

- Nice new measurement: <u>= production close to threshold</u>, Au-Au @ 6 A GeV (E895) [P.Chung et al.: nucl-ex/0302021]
- At this energy RQMD works fine!





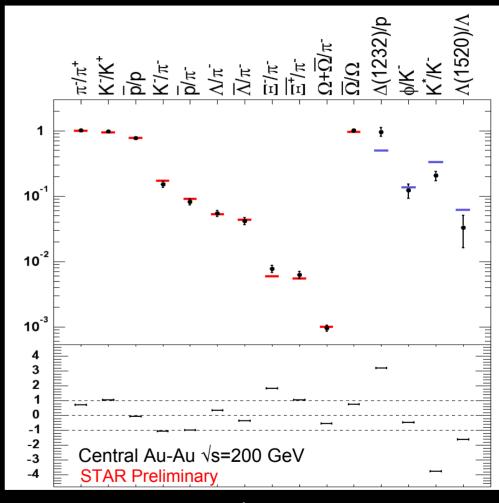
#### Thermal fits

 Thermal fits don't do too bad!

e.g.: brand new @ RHIC 200

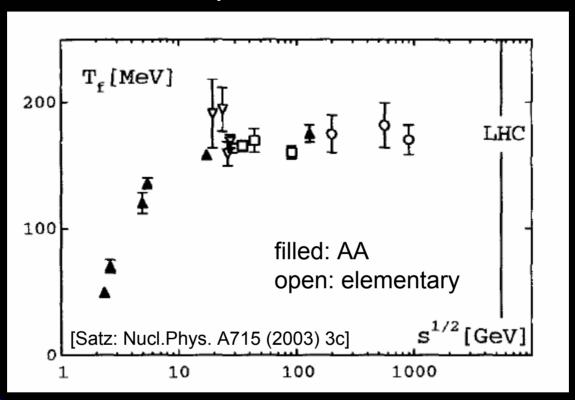
- $T_{ch} = 160 \pm 10 \text{ MeV}$
- $\bullet$   $\mu_B = 24 \pm 5 \text{ MeV}$

(STAR Preliminary)



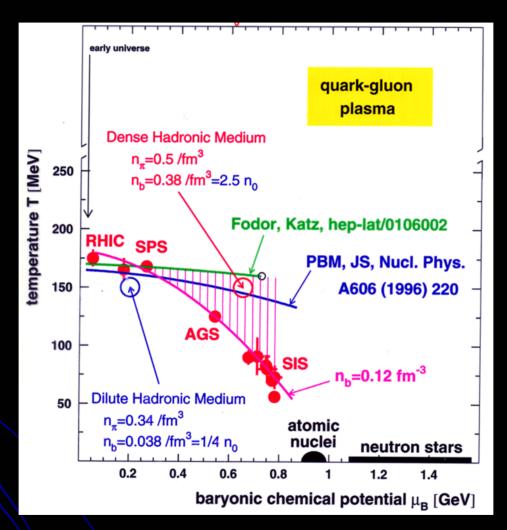
 relative particle abundances ~ as expected at thermodynamical equilibrium for grand-canonical system, even for rare, multi-strange particles

#### T systematics



- it looks like Hagedorn was right!
  - if the resonance mass spectrum grows exponentially (and this seems to be the case), there is a maximum possible temperature for a system of hadrons
  - indeed, we don't seem to be able to get a system of hadrons with a temperature beyond T<sub>max</sub> ~ 170 MeV!

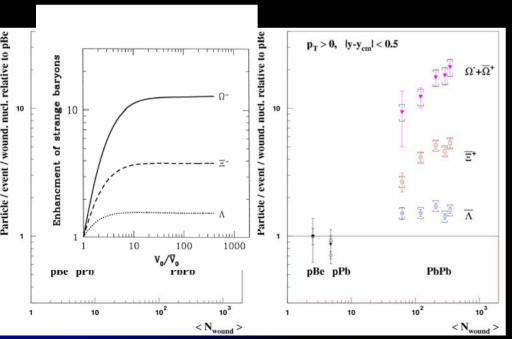
#### T vs $\mu_{B}$ systematics



 the extracted freeze-out points at SPS and RHIC lay very close to the predicted QGP phase boundary

#### Canonical vs Grand Canonical

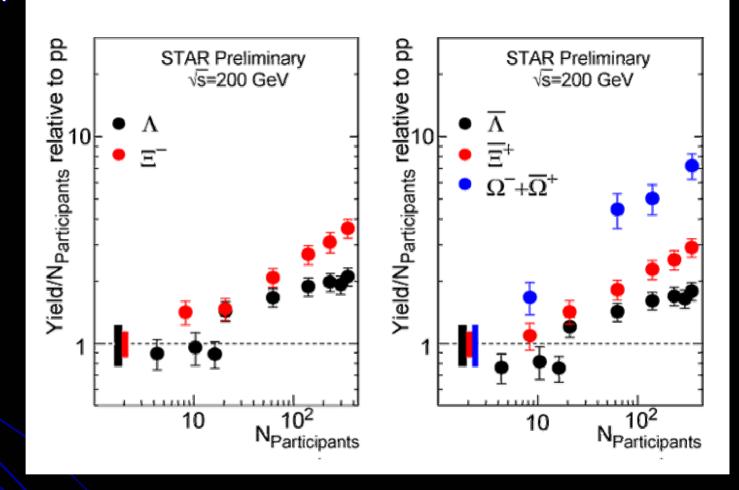
- Energy penalty to create a strange particle:
  - <u>Canonical</u>: computed taking into account also energy to create companion to ensure conservation of strangeness
  - Grand Canonical limit:
     just due to creation of particle itself. The rest of the system acts
     as a reservoir and "picks up the slack"



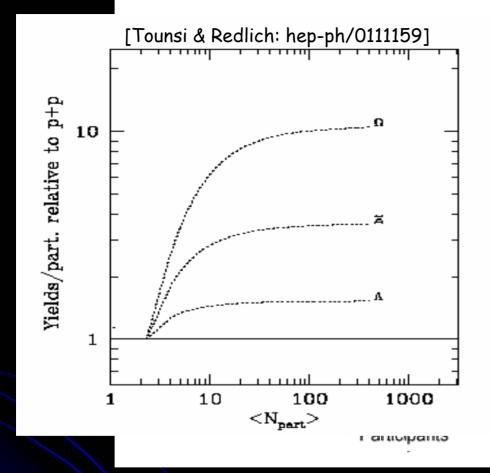
- removal of canonical suppression
  - increases with strangeness
  - ~ observed enhancements
  - detailed centrality dependence not reproduced (very crude modelling)

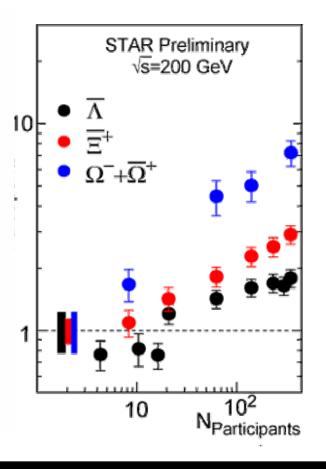
[Hamieh et al.: Phys. Lett. B486 (2000) 61]

#### • @ RHIC:



#### • @ RHIC:





- decrease of enhancements from SPS predicted
- order of magnitude of enhancements close
- again, model for V<sub>corr</sub> vs N<sub>part</sub> seems to be inadequate

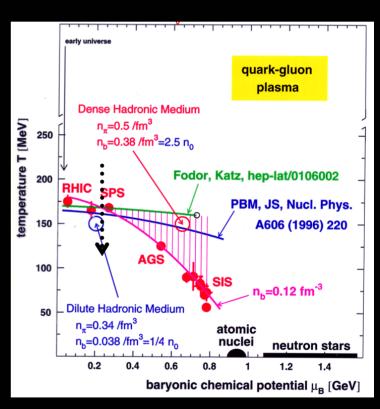
# Does this *explain* the observed enhancement pattern?

- a system <u>in eq.</u>, if it is large enough, is in GC eq., but being large in itself is not a sufficient condition for being GC!
  - if AA colls. were just a superposition of pp, they would have to be treated canonically all the same!
- the system must also know it is large...
  - it must know that an  $\Omega^+$  generated here can be compensated by, say, an  $\Omega^-$  on the other side of the fireball!
  - what counts is the correlation volume
- "Canonical Suppression is removed!"
  - an observation, not an explanation

# How is Canonical Suppression removed?

how does the system know it is large? how can information travel so quickly through the system?

- → not by conventional hadronic transport (no time!)
- natural if the system is coming back from deconfinement



## How is Canonical Suppression removed?

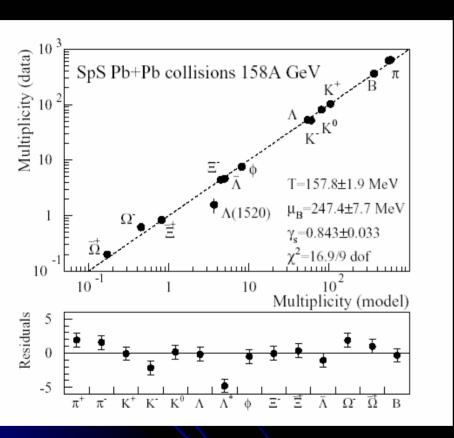
- how does the system know it is large?
  how can information travel so quickly through the system?
- → not by conventional hadronic transport (no time!)
- natural if the system is coming back from deconfinement
- two recent ideas:
  - quantum coherence in the fireball [Stock: hep-ph/0312039]
    - can it be developed into testable model?
    - what about pA?
  - multi-particle interactions [Braun-Munzinger et al.: nucl-th/0311005]
    - only important close to phase boundary
    - if true, we are actually *measuring* Tc
    - microscopic mechanism to restore abundances to hadronic eq.

### An equilibrium pre-diction

- mid-rapidity multi-strange/∧ @ 40 GeV SPS
  - NA57 vs [Becattini: private communication]
    - $\gamma_s = 1$ , T from  $4\pi$  fit to NA49 40 GeV [Becattini et al.: hep-ph/0310049] • using  $\pi^{\pm}$ ,  $K^{\pm}$ ,  $\varphi$ ,  $\Lambda$  and  $\overline{\Lambda}$
    - $\Lambda$  and  $\Lambda$  dN/dy from NA57 fix  $\mu_B$  and normalization
    - Prediction of  $\Xi^- \Xi^+ \Omega^-$  and  $\Omega^+ dN/dy$ 
      - total  $\chi^2$ /dof = 1.4 against values measured by NA57
  - → No parameter for multi-strange again, not bad...

# 

#### **\Lambda(1520)**



 Λ(1520) deviates from sistematics (~ 1/2)

e.g.: [Markert: J. Phys. G 28 (2002) 1753]

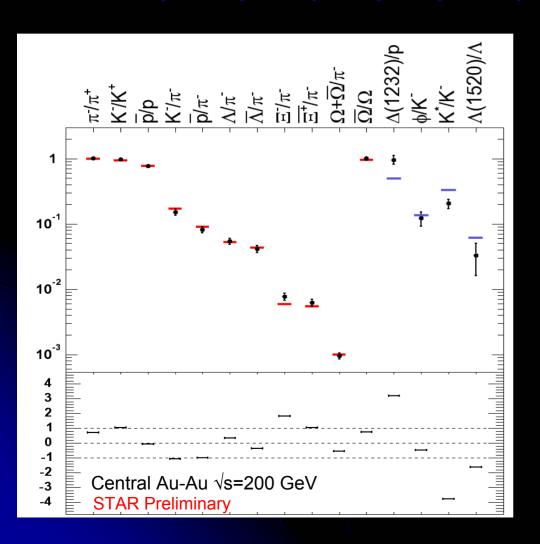
→ we look for in-medium effects

we take fits very seriously!

Fit to NA49 data
[Becattini et al.: hep-ph/0310049]

#### Resonances @ RHIC

new: Λ(1520), K\*(892), Δ(1232) in STAR at 200 GeV



- resonances show larger deviations from statistical fit
- rescattering vs regeneration?
  - see talk by P. Fachini

## Non equilibrium: $\gamma_s$

- ullet  $4\pi$  yield ratios sometimes preferred for thermal fits
  - to be safe in case strangeness ends up too far away in rapidity from where originally produced
  - (but can one assume global equilibrium with same T and same  $\mu_B$  at different rapidities?)
- in this case  $\gamma_s \sim 0.7$  0.8 must be introduced
  - strangeness undersaturation factor:
    - for each particle, a factor  $\gamma_s^{N(s+\overline{s})}$
    - e.g.: [Becattini et al.: hep-ph/0310049]
  - could it also mimic residual canonical suppression?

# Non equilibrium: Ys & Yq

see e.g.: [Rafelski & Letessier: hep-ph/0309030]

- $\gamma_q$  controls overall abundance of  $q\bar{q}$
- old  $\gamma_s$  becomes  $\gamma_s/\gamma_q$
- w.r.t. previous, additional factor  $\gamma_q^B$  for baryons
- fits get beautiful e.g.: @ RHIC →
   [see also Becattini et al.: hep-ph/0310049]
   (but additional parameter)
- @ RHIC 130:

$$\gamma_q = 1.6 \pm 0.2$$
  $\gamma_s / \gamma_q = 1.3 \pm 0.1$ 

 strangeness actually enhanced above hadronic equilibrium values?

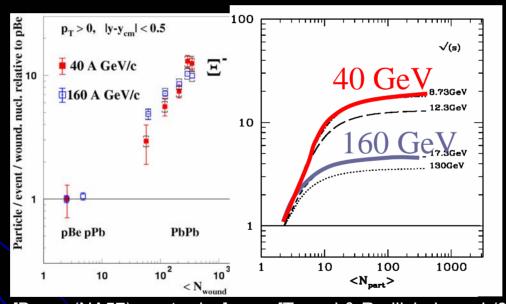
ratio	RHIC-130	non-eq fit	$\chi^2$			
$\pi^+/p$	$9.5 \pm 1.4$	9.07	1.15			
$\pi^-/p$	$13.4 \pm 0.9$	13.15	0.08			
$p/h^-$	_	0.0459	-			
$\Lambda_c/h^-$	$0.059 \pm 0.004$	0.0509	4.11			
$\Lambda_c/h^-$	$0.042 \pm 0.004$	0.0379	1.04			
$\Xi_c^-/h^-$	$0.0079 \pm 0.0012$	0.00805	0.01			
$\overline{\Xi_c^-}/h^-$	$0.0066 \pm 0.001$	0.00645	0.02			
$\Omega/h^-$	$(12 \pm 5)10^{-4}$	$13.210^{-4}$	0.06			
$(\overline{\Omega} + \Omega)/h^-$	$(22 \pm 6.5)10^{-4}$	$24.810^{-4}$	0.19			
$\Lambda_c/p$	$0.90 \pm 0.12$	0.747	1.63			
$\overline{\Lambda_c}/\overline{p}$	$0.93 \pm 0.19$	0.826	0.30			
$\Xi^{-}/\Lambda$	$0.193 \pm 0.03$	0.189	0.02			
$\overline{\Xi^-}/\overline{\Lambda}$	$0.219 \pm 0.035$	0.207	0.12			
$\Omega/\Xi^-$	_	0.164	_			
$\overline{\Omega}/\overline{\Xi^{-}}$	_	0.180	_			
$(\overline{\Omega} + \Omega)/(\overline{\Xi} + \Xi)$	$0.150 \pm 0.04$	0.171	0.28			
$\overline{p}/p$	$0.71 \pm 0.06$	0.674	0.36			
$\Lambda_c/\Lambda_c$	$0.71 \pm 0.04$	0.745	0.78			
₹/≘	$0.83 \pm 0.08$	0.801	0.13			
$\overline{\Omega}/\Omega$	$0.95\pm0.1$	0.878	0.51			
$\dot{K}^{+}/\pi^{+}$	$0.17 \pm 0.02$	0.195	1.59			
$K^-/\pi^-$	$0.15 \pm 0.02$	0.180	2.28			
$K^-/K^+$	$0.87 \pm 0.07$	0.923	0.57			
$K^{*0}/K^{-}$	$0.26 \pm 0.08$	0.231	0.13			
$\phi/h^{-}$	$0.02 \pm 0.002$	0.0212	0.37			
$\phi/K^-$	$0.15 \pm 0.03$	0.148	0.00			
$\phi/K^{*0}$	$0.595 \pm 0.24$	0.639	0.03			

$\sqrt{s_{NN}}$ [GeV]	200	130	200	130
T [MeV]	$143 \pm 7$	$144 \pm 3$	$160 \pm 8$	$160 \pm 4$
$\mu_b  [{ m MeV}]$	$21.5\pm31$	$29.2 \pm 4.5$	$24.5\pm3$	$31.4 \pm 4.5$
$\mu_s  [{ m MeV}]$	$2.5 \pm 0.2$	$3.1\pm0.2$	$2.9 \pm 0.2$	$3.6\pm0.2$
$\mu_{ m S}  [{ m MeV}]$	$4.7 \pm 0.4$	$6.6 \pm 0.4$	$5.3 \pm 0.4$	$6.9 \pm 0.4$
$\gamma_q$	$1.6 \pm 0.3^{*}$	$1.6 \pm 0.2^*$	1*	1*
$\gamma_s/\gamma_q$	$1.2 \pm 0.15$	$1.3\pm0.1$	$1.0 \pm 0.1$	$1.13\pm0.06$
$\chi^2/dof$	2.9/6	15.8/24	4.5/7	32.2/25
$P_{ m true}$	90%+	95%+	65%	15%

- pentaquark production (if statistical) would be particularly sensitive to need for  $\gamma_q!$ 
  - additional  $\gamma_q^2$  factor for  $\Theta^+$  w.r.t. other models e.g.: [Letessier et al.: hep-ph/0310188, Becattini et al.: hep-ph/0310049]

## Hyperon enhancements at 40 GeV

- enhancements vs N<sub>wound</sub> steeper at 40 than at 160
- C to GC predicts a factor 4 5 larger Ξ⁻ enhancement at 40 GeV than at 160, not seen in the data

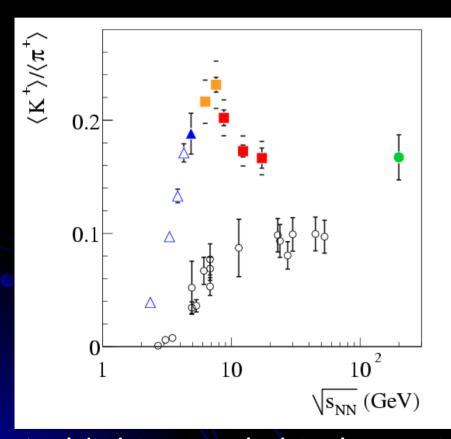


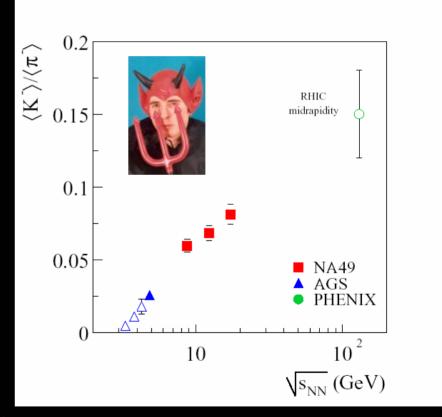
[Bruno (NA57): yesterday] [Tounsi & Redlich: hep-ph/0111159]

- perhaps yields trying to shoot up but don't go all the way?
  - on the other hand, equilibrium fit not too bad... [Becattini: priv. comm]

#### Marek's Horn

- Change of behaviour of  $K^+/\pi^+$  in A-A around 30 GeV
  - $K^-/\pi^-$  has a smooth behaviour with energy

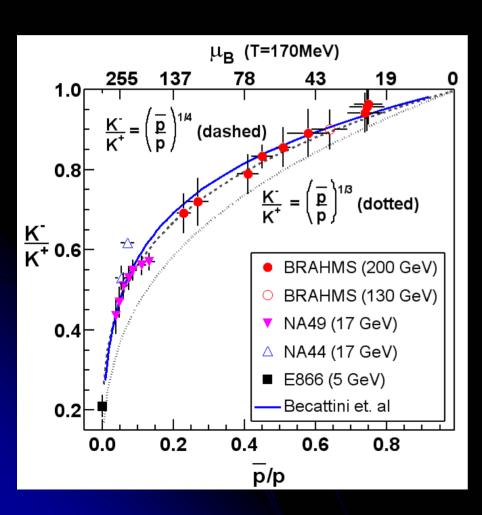


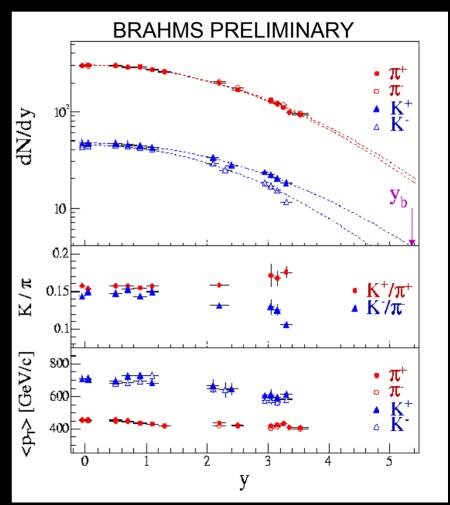


Could this signal chiral transition?

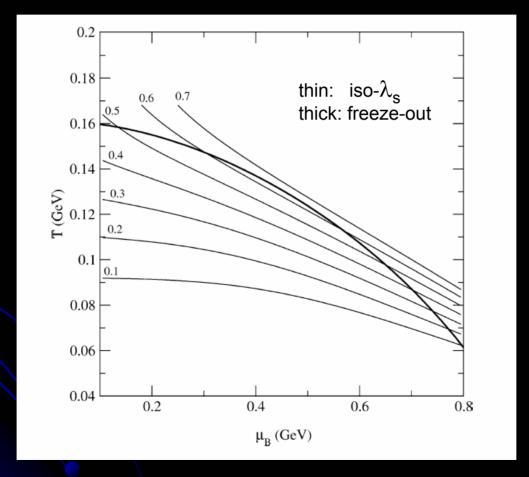
### Brahms: K± vs y

caution! K<sup>±</sup> sensitive to baryon density...



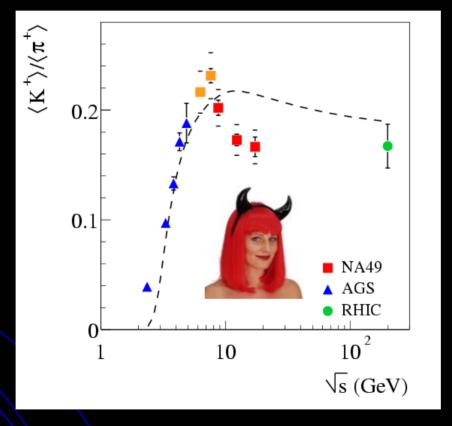


#### • ... and $\lambda_s$ [ = $2 < s\overline{s} > /(\langle u\overline{u} \rangle + \langle d\overline{d} \rangle)$ ] max. around 30 GeV:



[Braun-Munzinger et al.: NPA697(2002)902]

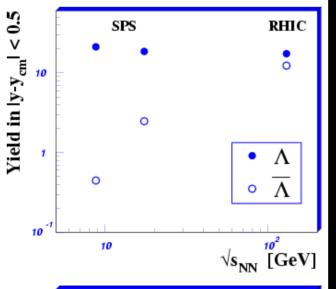
• so one expects a broad maximum in  $K^+/\pi^+$  about there...

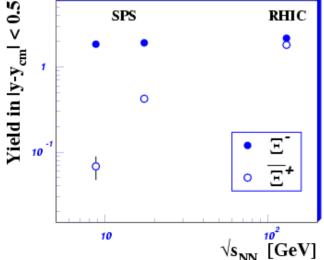


- but nothing so sharp!
  - → it looks like we have some anomaly...

## Absolute values of hyperon dN/dy

see talk by D.Elia, Friday, P2

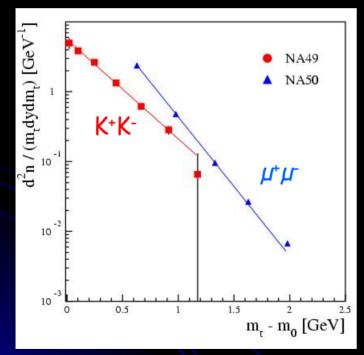




- absolute values of  $\Lambda$  and  $\Xi^+$  dN/dy go up with  $\sqrt{s}$ ...
- absolute values of A and Ξ⁻ dN/dy
   constant from √s<sub>NN</sub> = 8.8 to 130 GeV!
- Statistical fits: T and  $\mu_B$  only sensitive to particle ratios
  - overall normalization of yields controlled by fitted V parameter
- T,  $\mu_B$  and V all vary with energy, but in such a way as to ensure  $\Lambda$ ,  $\Xi^-$  yields stays constant ...
  - there must be a simple reason ...

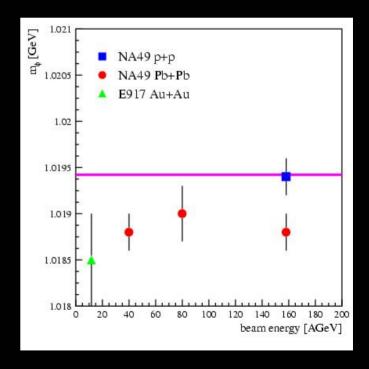
#### The mysterious **P**

- ullet fundamental to understand  $\gamma_s$  vs C suppression
  - $\gamma_s^2$ , but no C suppression
- the NA49 NA50 discrepancy is still with us...



 $\rightarrow$  NA60 should provide high stats, low p<sub>T</sub>  $\Phi \rightarrow \mu\mu$ 

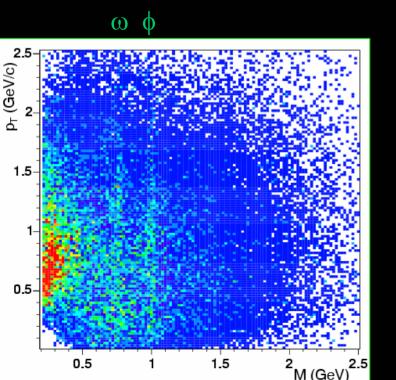
 NA49: small deviation from PDG mass seen in all three Pb+Pb data sets, but not in pp

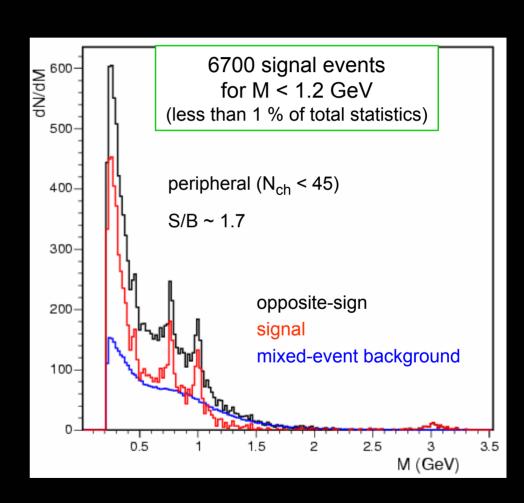


see [NA49 (Friese): SQM 03]

## First & signal from NA60

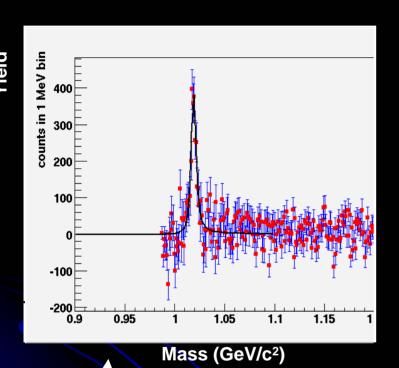
- peripheral In-In
- expected > 100k  $\Phi \rightarrow \mu^{+}\mu^{-}$
- expected mass resolution:~ 20 MeV
- also looking at  $\Phi \rightarrow K^+K^-$





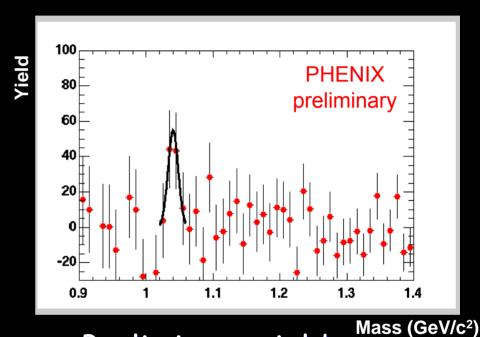
#### **Penix**

•  $\Phi \rightarrow K^+K^-$  (min bias 200 GeV)



Mass and width agree within errors with PDG values

•  $\Phi \rightarrow e^+e^-$  (min bias 200 GeV)



Preliminary yields:

$$\phi \rightarrow e^+ e^- : \frac{dN}{dy} = 5.4 \pm 2.5 (stat)_{-2.8}^{+3.4} (sys)$$

$$\phi \rightarrow K^+K^-: \frac{dN}{dy} = 2.01 \pm 0.22 \,(stat)_{-0.52}^{+1.01} \,(sys)$$

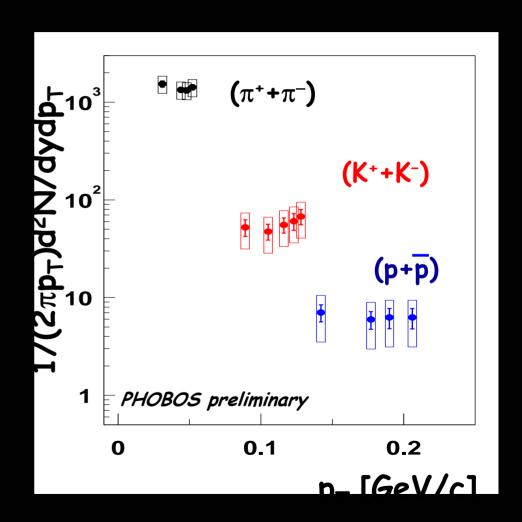
# Extras...

some results on spectra that I could not resist showing...

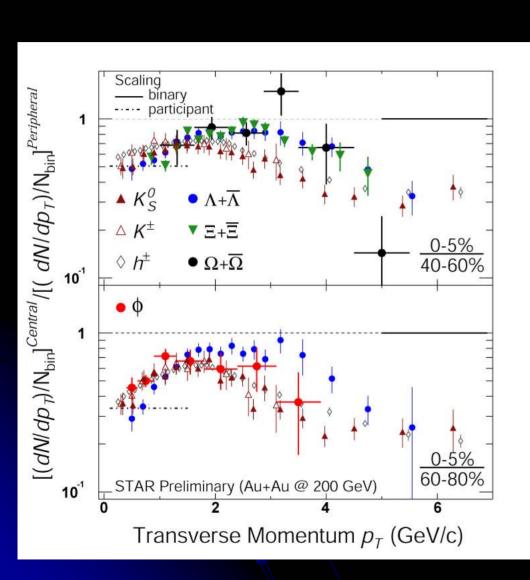
#### Very low $p_T \pi$ , K, p from Phobos!

- tracks characterized by range and energy loss in Si layers
  - → particle id.
  - → momentum within 5%

 unique constraints for transport, hydro, thermal models

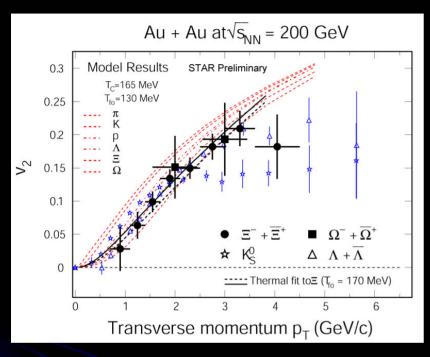


## Rcp for multistrange!



- $\Phi$  seems to follow  $K^0_s$  more than  $\Lambda$ 
  - → meson-baryon, not mass?

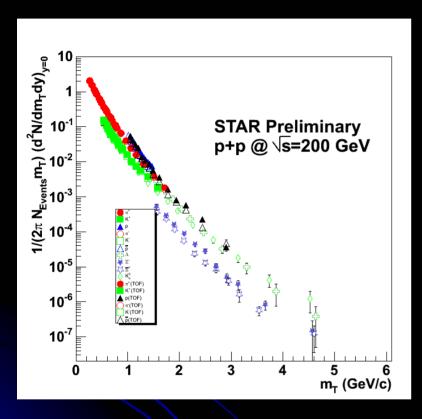
## v2 for multistrange!



 scaling seems to work (not for pions...)
 → parton recombination? 0.15  $\vdash$  n=2  $\diamond \pi^+ + \pi^- PHENIX$ •  $\Lambda$ + $\overline{\Lambda}$  STAR △ K++K- PHENIX +  $\Xi$ + $\overline{\Xi}$  STAR  $^{\blacktriangle}$   $K_{S}^{0}$  STAR 0.1 0.05 **STAR Preliminary** Transverse Momentum p<sub>⊤</sub>/n (GeV/c)

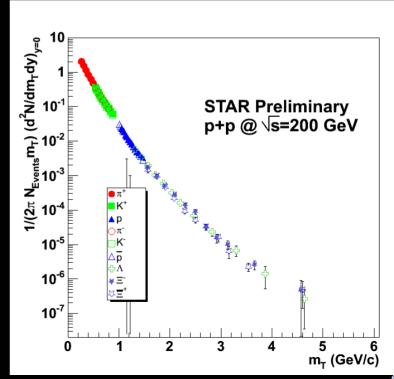
scaled with n(quarks)

# m<sub>T</sub> scaling in pp...



not absolute...

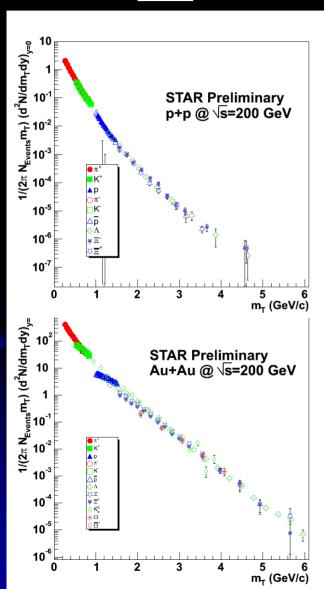
...but if one rescales:



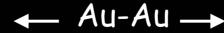
#### ... but not in Au-Au

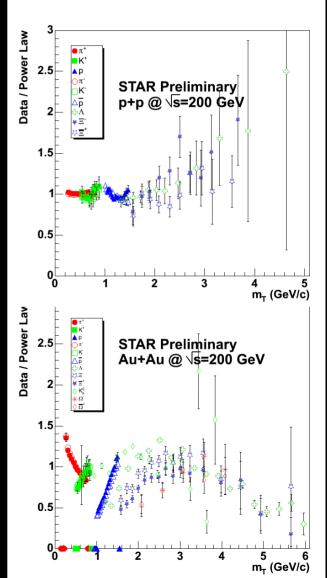
data

#### data / power law









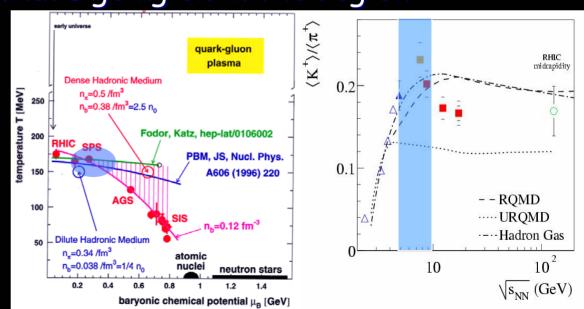
#### Conclusions

- Evidence for high degree of chemical equilibration
  - even for rare, multistrange particles
  - → QGP prediction!
  - → freeze-out curve traces expected phase boundary at SPS - RHIC
  - →hadronic transport does not seem to be able to do it (no time!)
- RHIC opening up exciting new windows for strangeness!
  - elliptic flow, high p<sub>T</sub>, high statistics reference data (pp, dAu)
- Strangeness perhaps enhanced above hadronic equilibrium!
  - stand-alone smoking gun?

- we must understand (non-)equilibrium
  - we need highest possible precision data on yields and ratios (RHIC!)
  - (we need theory to provide agreed statistical approach)
  - we need to understand the Φ
    - precision data from NA60 and Phenix
- we must understand what's going on in the region

√s<sub>NN</sub> = 5 - 10 GeV

- low E SPS?
- SIS 200 @ GSI!



#### Credits

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